VOLCANICA

Estimation and uncertainty quantification of magma interaction times using statistical emulation

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ABSTRACT

Evolution of volcanic plumbing systems towards eruptions of different styles and sizes largely depends on processes at crustal depths that are outside our observational capabilities. These processes can be modelled and the outputs of the simulations can be compared with the chemistry of the erupted products, geophysical and geodetic data to retrieve information on the architecture of the plumbing system and the processes leading to eruption. The interaction between magmas with different physical and chemical properties often precedes volcanic eruptions. Thus, sophisticated numerical models have been developed that describe in detail the dynamics of interacting magmas, specifically aimed at evaluating pre-eruptive magma mingling and mixing timescales. However, our ability to explore the parameter space in order to match petrological and geophysical observations is limited by the extremely high computational costs of these multiphase, multicomponent computational fluid dynamics simulations. To overcome these limitations, we present a statistical emulator that is able to reproduce the numerical simulations results providing the temporal evolution of the distribution of magma chemistry as a function of a set of input parameters such as magma densities and reservoir shapes. The whole rock composition of volcanic rocks is one of the most common measurable parameter collected for eruptions. The statistical emulator can be used to invert the observed distribution of whole rock chemistry to determine the duration of interaction between magmas preceding an eruption and identify the best matching input parameters of the numerical model. Importantly, the statistical emulator intrinsically includes error propagation, thus providing confidence intervals on predicted interaction timescales on the base of the intrinsic uncertainty of the input parameters of the numerical simulations.

KEYWORDS: Volcanology; Petrology; Magma Dynamics; Statistical Emulation.

1 INTRODUCTION

Magmatic processes preceding volcanic eruptions occur at depths of kilometres to tens of kilometres and are therefore not directly accessible: the only direct information about preeruptive magma storage conditions is provided by the volcanic rock record. Thus, textural [\[Cashman and Marsh 1988;](#page-11-0) [Hig](#page-12-0)[gins 2000;](#page-12-0) [Zellmer et al. 2020\]](#page-14-0), rheological [\[Lejeune and Richet](#page-13-0) [1995;](#page-13-0) [Hess and Dingwell 1996;](#page-12-1) [Lavallée et al. 2007;](#page-13-1) [Pistone et](#page-14-1) [al. 2012;](#page-14-1) [Heap et al. 2020\]](#page-12-2), and chemical [\[Marsh 1981;](#page-13-2) [Blundy](#page-11-1) [and Cashman 2001;](#page-11-1) [Davidson et al. 2007;](#page-11-2) [Putirka 2008;](#page-14-2) [Ed](#page-12-3)[monds et al. 2010;](#page-12-3) [Humphreys et al. 2012;](#page-12-4) [Cooper and Kent](#page-11-3) [2014;](#page-11-3) [Neave et al. 2014;](#page-13-3) [Cassidy et al. 2016;](#page-11-4) [Stock et al. 2018;](#page-14-3) [Edmonds et al. 2019;](#page-12-5) [Maclennan 2019;](#page-13-4) [Zellmer et al. 2020\]](#page-14-0) characterisation of volcanic rocks is central to our understanding of the processes leading to volcanic eruptions. However, to retrieve quantitative information on pre-eruptive processes from the analysis of the rock record, it is fundamental frst to compare the chemistry of minerals, rocks, and glasses with those produced experimentally [\[Nimis 1995;](#page-13-5) [Putirka 2008;](#page-14-2) [Pe](#page-14-4)[trelli et al. 2020\]](#page-14-4) and then to compare the analyses with the results of experiments or numerical modelling [e.g. [Annen et](#page-11-5) [al. 2006;](#page-11-5) [Caricchi et al. 2014;](#page-11-6) [Perugini et al. 2015;](#page-14-5) [Jackson](#page-12-6)

[et al. 2018\]](#page-12-6). We will focus here on the comparison between volcanic rock record and numerical modelling results.

Physical modelling is a fundamental tool to link observable quantities at the Earth surface and magmatic processes at depth [e.g. [Bagagli et al. 2017;](#page-11-7) [La Spina et al. 2022\]](#page-13-6). Modelling efforts have been directed both at tackling the complex thermochemistry of multiphase magmas [e.g. [Gualda et al. 2012;](#page-12-7) [Bohrson et al. 2014;](#page-11-8) [Keller and Suckale 2019;](#page-13-7) [Rummel et al.](#page-14-6) [2020\]](#page-14-6), in order to be able to replicate observed geochemical patterns in erupted products (crystals, melt inclusions [e.g. [Caricchi et al. 2018;](#page-11-9) [Wei et al. 2022\]](#page-14-7)) as well as emitted gas chemistry and fuxes [e.g. [Chiodini et al. 2012\]](#page-11-10), and at replicating observed geophysical signals such as ground deformation, gravity anomaly, and seismicity recorded at active volcanoes [e.g. [Poland and de Zeeuw-van Dalfsen 2021;](#page-14-8) [Matoza](#page-13-8) [and Roman 2022\]](#page-13-8). Inclusion of all relevant physics is often impossible, given the wide variety of length and time scales that characterise volcanic systems at large; therefore, diferent approximations are typically used to target specifc questions: homogeneous reservoirs are often assumed in order to model ground deformation [e.g. [Mogi 1958;](#page-13-9) [Zhong et al. 2019\]](#page-14-9); magma motion is usually neglected when modelling crystal growth patterns [e.g. [Iovine et al. 2017;](#page-12-8) [Petrone et al. 2018\]](#page-14-10). Necessity of simplifying assumptions is also motivated by the very high computational costs of running large-scale, detailed

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simulations of the overall magmatic system evolution [\[Garg](#page-12-9)] [and Papale 2022\]](#page-12-9). This also largely limits our ability to explore the wide range of initial conditions and input parameters that characterise magmatic systems, such as temperatures and pressures/depths of melt reservoirs, as well as their size, shape, and melt chemistry, and volatile content. Moreover, as properties of magma reservoirs at depth cannot be measured, models' initial conditions are typically characterised by large uncertainties which are—in turn—difficult to define.

Due to the necessary assumptions and uncertainties associated with the range of some of the input parameters of physical models, it is often impossible to quantitatively assess the impact of initial assumptions on the fnal results of the models. This would require the realisation of millions of simulations to systematically investigate the full range of input parameters. Additionally, the use of complex physical models during volcanic unrest can be unpractical as the realisation of simulations can be time-intensive and the integration of continuous streams of data not straightforward [\[Selva et al. 2012\]](#page-14-11).

This is where the combination of physical modelling and statistical emulation becomes essential. Using the outputs of simulated physical states and the corresponding values for the theoretical parameters—i.e. simulated data produced by a simulator given a set of input parameters—a Statistical Emulator (SE) allows one to predict the physical model output given a specifc input, along with an associated measure of the uncertainty on this prediction. More precisely, an SE is a probabilistic model that links any output of a simulator, often summary statistics of the simulated data, produced by a theoretical physical model (i.e. the simulator), to the parameter values used to produce the simulation results. Once trained on a series of simulated data, the SE can be used on observed data to calibrate the possible values of some of the input parameters of the underlying physical process. Statistical emulation is a very popular method in environmental sciences. While [Sacks](#page-14-12) [et al.](#page-14-12) [\[1989\]](#page-14-12), [Currin et al.](#page-11-11) [\[1991\]](#page-11-11), [Kennedy and O'Hagan](#page-13-10) [\[2001\]](#page-13-10), [Santner et al.](#page-14-13) [\[2003\]](#page-14-13), and [Higdon et al.](#page-12-10) [\[2008\]](#page-12-10) use a Gaussian process to model the uncertainty, this requires long computational times and a large number of SEs; in practice, a (multivariate) regression model is often preferred (see e.g. [Salter](#page-14-14) [and Williamson](#page-14-14) [\[2016\]](#page-14-14) for a discussion and references). Notably, diferent SEs that rely on various statistical assumptions are regularly proposed, and they depend both on the physical model and the available data; see e.g. [Mahmood et al.](#page-13-11) [\[2015\]](#page-13-11), [Guillas et al.](#page-12-11) [\[2018\]](#page-12-11), and [Yang et al.](#page-14-15) [\[2020\]](#page-14-15) to cite just a few.

In this paper we present results from a numerical model of magma mingling, tailored for the Campi Flegrei magmatic system (Italy) [\[Montagna et al. 2022\]](#page-13-12). We show how silicate melt composition evolves as two parental magmas interact in a relatively shallow reservoir at 3 km depth, approaching compositional homogeneity [\[Montagna and Papale 2018\]](#page-13-13), for diferent setups characterised by varying reservoir shape and magmatic volatile contents. An SE is then trained on these results; we present the high accuracy it achieves in reproducing simulated results, even at large down-sampling, and the error quantifcation associated with this method.

2 MAGMA DYNAMICS MODELS

The very existence of magma within the Earth's crust requires prolonged input of mass and energy (i.e. heat) from the mantle source [\[Glazner et al. 2004;](#page-12-12) [Annen et al. 2006;](#page-11-5) [Karakas et al.](#page-12-13) [2017\]](#page-12-13). Thus, the progressive assembly of volcanic plumbing systems naturally results in the interaction between resident, and generally colder, magma and hot, generally more primitive magma rising through the crust. Abundant feld evidence exists for this process, both in the plutonic and volcanic rock records [\[Blundy and Sparks 1992;](#page-11-12) [Perugini and Poli 2005;](#page-14-16) [Mor](#page-13-14)[gavi et al. 2016;](#page-13-14) [Ridolf et al. 2016;](#page-14-17) [Morgavi et al. 2017\]](#page-13-15), and in some cases such interaction can culminate in a volcanic eruption [\[Caricchi et al. 2021\]](#page-11-13). Examples of magma input triggering volcanic activity can be identifed at Campi Flegrei, Santorini (Greece), Yellowstone Caldera (USA), and Long Valley Caldera (USA) [\[Wark et al. 2007;](#page-14-18) [Druitt et al. 2012;](#page-12-14) [Morgavi et al. 2016;](#page-13-14) [Forni et al. 2018\]](#page-12-15), and approaches have been proposed to determine the duration of magma interaction preceding an eruption [\[Perugini et al. 2015\]](#page-14-5). Here, we simulate the dynamical evolution of a melt-dominated magmatic reservoir subject to the injection of deeply-sourced, more primitive, more volatilerich (buoyant) magma [\[Montagna et al. 2015;](#page-13-16) [2022\]](#page-13-12), focusing on the quantifcation of the time scales of magmatic convection and mingling.

2.1 Physical model setup

Numerical simulations of magma chamber replenishment were performed using the fnite-element C++ code GALES [\[Longo et al. 2012a;](#page-13-17) [Garg et al. 2018a;](#page-12-16) [b\]](#page-12-17). The physicomathematical model describes the space-time dynamics of a multicomponent mixture consisting of silicate melt in thermodynamical equilibrium with a gas phase at the local conditions of pressure, temperature and composition [\[Longo et al. 2012b;](#page-13-18) [Papale et al. 2017;](#page-13-19) [Garg et al. 2019\]](#page-12-18). The conservation equations for the mass of single components and momentum of the whole mixture are solved together with the gas-liquid thermodynamic equilibrium model [\[Papale et al. 2006\]](#page-14-19) and the constitutive equations for multiphase, multicomponent mixture properties density and viscosity [\[Reid et al. 1977;](#page-14-20) [Ishii and Zu](#page-12-19)[ber 1979;](#page-12-19) [Lange 1997;](#page-13-20) [Giordano et al. 2008\]](#page-12-20).

We explore buoyant magma mixing and its timescales based on the archetypal case of the Campi Flegrei volcanic system, where occurrences of interacting magmas have been widely testifed [\[Tonarini et al. 2009;](#page-14-21) [Arienzo et al. 2010;](#page-11-14) [Four](#page-12-21)[mentraux et al. 2012;](#page-12-21) [Forni et al. 2018\]](#page-12-15). We model the injection of CO2-rich, shoshonitic magma coming from a deep reservoir into a shallower, much smaller one, containing more evolved and partially degassed phonolitic magma [\[Mangiacapra et al.](#page-13-21) [2008;](#page-13-21) [Arienzo et al. 2009;](#page-11-15) [Di Renzo et al. 2011\]](#page-12-22). The interaction dynamics is solely driven by buoyancy, stemming from the initial unstable density contrast between the shoshonitic and phonolitic end-members [\[Longo et al. 2023\]](#page-13-22).

Diferent geometries and volatile contents of the shallow chamber were considered to account for a range of possible conditions at Campi Flegrei. Specifcally, the shallow reservoir is oblate (Simulations 1 and 4 in [Table 1\)](#page-2-0), circular (Simulation 3 in [Table 1\)](#page-2-0), or prolate (Simulations 2 and 5 in [Table 1\)](#page-2-0); the surface area is kept constant. Total (dissolved and excess fuid) volatile contents in the phonolitic magma are 0.3 wt.% $CO₂$ and 2.5 wt.% $H₂O$ for Simulations 1, 2, and 3 in [Table 1,](#page-2-0) while for Simulations 4 and 5 they are set to 0.1 wt.% $CO₂$ and 1 wt.% $H₂O$ [\[Mangiacapra et al. 2008\]](#page-13-21). [Figure 1](#page-3-0) shows the simulated domain and summarises the conditions for the numerical simulations; the deep chamber contains a shoshonitic magma with 2 wt.% H_2O and 1 wt.% CO_2 .

The different settings in [Figure 1](#page-3-0) can be represented by two meaningful quantities, the initial density diference at the interface among the two end-member magmas (shallow phonolite and deep shoshonite) $\Delta \rho$, and the aspect ratio of the shallow reservoir a . [Table 1](#page-2-0) shows the values of these parameters for the five simulated scenarios. The initial density contrast is due to diferent volatile contents in the shallow reservoir, as well as diferent phase partitioning in volatiles. The latter depends on pressure, in turn a function of the depth of the interface which changes according to reservoir shape.

Table 1: Values of the input parameters for each simulated scenario.

	Aspect ratio a.	Density difference $\Delta \rho$ [kg m ⁻³]
Simulation 1	0.5	35
Simulation 2	20	20
Simulation 3	1.0	30
Simulation 4	0.5	160
Simulation 5	20	130

2.2 Magma mingling

As the two magmas come into contact at time 0, a gravitational Rayleigh-Taylor instability [\[Chandrasekhar 2013\]](#page-11-16) develops at the interface because of the density diference Δρ between the denser magma in the shallow reservoir and the volatile-rich and less dense injected magma. [Figure 2](#page-4-0) shows the space-time evolution of the composition within the shallower, initially phonolitic reservoir where shoshonitic magma is injected.

Plumes of light magma rise into the shallow reservoir and trigger convective patterns enhancing magma mixing, while a portion of the degassed phonolitic magma initially hosted in the shallow reservoir sinks through the feeding duke. With time, less dense shoshonitic magma tends to accumulate at the top of the system, and a density-stratifed magma chamber develops, more so when convection is more afficient, i.e. at small aspect ratios and large initial density contrasts [\(Figure 2\)](#page-4-0) [\[Montagna et al. 2015\]](#page-13-16). The space-time evolution of composition within the reservoir depends both on its aspect ratio a and on the initial density contrast $\Delta \rho$ [\(Figure 2\)](#page-4-0), but in all cases it testifes for the progressive homogenisation of magma chemistry [\[Montagna et al. 2015\]](#page-13-16). The time duration and efficiency of the interaction only depend on the initial conditions, i.e. $\Delta \rho$ and *a* [\[Papale et al. 2017\]](#page-13-19).

During the magmatic interaction described above, an eruption can occur, sampling the magma from the reservoir. Here we assess the capacity of the distribution of erupted magma chemistry to return information on the duration of interaction

preceding the eruption. This sort of information is essential for determining, from the study of past eruptions triggered by the interaction between magmas, the amount of time available between potential monitoring signs of magma input into a relatively shallow reservoir and an eruption [\[Bagagli et al. 2017\]](#page-11-7). Instead of performing a large number of numerical simulations, we resort to statistical emulation. We use the distribution of magma chemistry as a function of time obtained from the numerical simulations to fit the SE. The final target of this approach is to build a model which, on the basis of the distribution of erupted magma chemistry, provides the timescales of magmatic interaction that preceded the eruption, and the associated uncertainties.

Magmatic composition in the model is defned as the proportion of one end-member (shoshonite) with respect to the total, thus it varies between 0 and 1. End-member compositions are in turn defined by the ten major oxides $(SiO₂, TiO₂)$, AlO2, Fe2O3, FeO, MnO, MgO, CaO, Na2O, K2O [\[Montagna](#page-13-16) [et al. 2015\]](#page-13-16)), plus volatiles (main components water and $CO₂$ only). The timescales of the dynamics described by these simulations are very short compared to typical difusive timescales for oxides in magmas at our resolution (minimum grid size is ∼1 m [\[Zhang and Gan 2022\]](#page-14-22)), therefore the approximation of mingling holds and the modelled compositions compare with whole rock datasets. The oxides compositions of the two endmember magmatic components, as well as their volatile contents partitioned between liquid and exsolved phases, are in turn used in the model to evaluate volatile phase partitioning, and density and viscosity of the fuid, thus they have a signifcant impact on the evolution of the system [\[Longo et al. 2012a;](#page-13-17) [Montagna et al. 2015;](#page-13-16) [Longo et al. 2023\]](#page-13-22).

3 A STATISTICAL EMULATOR

An SE is a probabilistic model that links the output of a simulator ($\mathbf{D} \in \mathcal{D} \subset \mathbb{R}^{K_1}$), or a suitable function of the simulated data $(Y = Y(D) \in \mathcal{Y} \subseteq \mathbb{R}^{K_2}$, with $K_2 \ll K_1$, produced by a theoretical (complex) model (simulator), to the parameter values ($\mathbf{x} \in \mathcal{X} \subseteq \mathbb{R}^{K_3}$) used to produce the simulation results. Very generally, one considers theoretical models that are complex enough so that simulating one possible outcome, i.e. the outcome obtained with one (set of) value(s) for the parameters governing the theoretical model (i.e. $\mathbf{x}_i \in \mathcal{X} \subseteq \mathbb{R}^{K_3}$), is computationally very demanding. In these cases, the simulator cannot simulate information for all sets of possible parameters \mathbf{x}_i , but instead can simulate from a finite sample, say \mathbf{x}_i for $i = 1, \ldots, m$, of carefully chosen combinations of values for the \mathbf{x}_i 's.

In the context of time-dependent outputs, the set of parameters \mathbf{x}_i also contains the time points $t \in \{t_1, \ldots, t_{n_i}\}\,$, that can be conveniently separated into n_i subsets \mathbf{x}_{it} , i.e. one for each time point. The function of the simulated data, e.g. J summary statistics, can be denoted as Y_{ijt} , and in full generality, an SE postulates a relationship of the form

$$
Y_{ijt} = f_j(\mathbf{x}_{it}, \mathbf{\Theta}_j) + \varepsilon_{ijt},
$$
\n(1)

where $\mathbf{\Theta}_j$, for $j = 1, \ldots, J$, are the SE's parameters that are allowed to difer according to the diferent summary statistics. The emulators f_i can be the same functions for all

Figure 1: Summary of the five simulated settings. The upper chamber is either elliptical, with semi-axes of 200 and 400 m, or circular, with radius of 283 m to keep the same surface area. The dyke, which at time 0 hosts shoshonite like the deep chamber, is 20 m wide. The deep reservoir is elliptical, with semi-axes of 4.0 and 0.5 km. [A] full simulated domain, with prolate shallow reservoir; grey rectangle indicates the region magnified in [B], [C] and [D]. [B] Simulations 1 and 4 with oblate shallow reservoir. [C] Simulation 3 with circular shallow reservoir. [D] Simulations 2 and 5 with prolate shallow reservoir.

j, but they differ in the values of the parameters θ_i . The stochastic term ε_{ijt} allows to model what is not captured by the emulator and is defned after [Equation 2.](#page-3-1) Thus, in vector form, we can consider $Y_{ij} = f_j(x_i, \theta_j) + \varepsilon_{ij}$ $\left[f_j(\mathbf{x}_{it_1}, \boldsymbol{\theta}_j), \ldots, f_j(\mathbf{x}_{it_{n_i}}, \boldsymbol{\theta}_j)\right]^T + \boldsymbol{\varepsilon}_{ij}$, or, in a more compact form:

$$
\mathbf{Y}_i = f(\mathbf{x}_i, \boldsymbol{\theta}) + \boldsymbol{\varepsilon}_i
$$
\n
$$
= \left[f_1(\mathbf{x}_{it_1}, \boldsymbol{\theta}_1), \dots, f_1(\mathbf{x}_{it_{n_i}}, \boldsymbol{\theta}_1), \dots, \dots, f_J(\mathbf{x}_{it_{n_i}}, \boldsymbol{\theta}_J) \right]^T + \boldsymbol{\varepsilon}_i. \quad (2)
$$

One also supposes, very generally, that ε_i = $[\varepsilon_{ijt}]_{j=1,...,Jt=t_1,...,t_{n_i}}$ is a stochastic process such that $\epsilon_i \sim \mathcal{F}\left\{0, \Sigma\right\}$, where \mathcal{F} denotes some probability distribution in $\mathbb{R}^{J \times n_i}$ with mean 0 and covariance Σ , where $\Sigma_{kl} = \sigma_k^2$ for $k = l$ and its off-diagonal terms can be non-zero. The vector of parameters $\mathbf{\theta} = [\mathbf{\theta}_1^T]$ T_1 ... $\mathbf{\theta}_J^T$ $]^T \in \mathbf{\Theta} \subseteq \mathbb{R}^p$ is unknown, and needs to be estimated given a series of simulations $D = (D_1, \ldots, D_m)$ from which *J* summary statistics are computed at n_i time points to obtain the Y_i 's. The noise ε_i captures the approximation error of the link functions f_i in [Equation 2,](#page-3-1) which can be made as complex as possible, but will never match exactly the true underlying theoretical process. The noise is random, and we suppose that its (probability) distribution is fully specifed by its first two moments—i.e. for the *i*-th simulation, $\mathbb{E}[\epsilon_i] = 0$ and $cov(\varepsilon_i) = \Sigma$. Importantly, the components of Σ can be estimated to obtain an estimated covariance matrix $\hat{\Sigma}$, which can be used to compute confdence intervals for the predictions provided by the SE.

3.1 Application to magma dynamics

We consider here an SE for the magma dynamics model presented in [Section 2.](#page-1-0) The salient features of the SE, as well as the construction of its inputs, are highlighted in the flow chart of [Figure 3.](#page-5-0) The input parameters \mathbf{x}_{it} , for the $m = 5$ simulations described in [Table 1,](#page-2-0) are the aspect ratio of the

Figure 2: Space-time evolution of composition in the upper region of the simulated domain, for the different scenarios. Rows represent different times, columns represent different simulations.

shallow reservoir a_i , the initial density difference between the two end-member magmas $\Delta \rho_i$, and the interaction time $t \in \{t_1, \ldots, t_{n_i}\}\$, so that $\mathbf{x}_{it} = [a_i, \Delta \rho_i, t]^T$. The output from the simulator is D_i , an $r_i \times n_i$ matrix containing the silicate melt compositions expressed as the proportions of the two end-member magmas, which are measured across r_i locations of the shallow reservoir at n_i interaction times. From this simulator output, summary statistics are computed to reduce the output dimension, and we have chosen to consider $J = 5$ quantiles (namely, 0.05, 0.25, 0.50, 0.75, and 0.95) computed from the distribution of the compositions over the r_i locations, with $J \ll r_i$. Namely, Y_{ijt} in [Equation 1](#page-2-1) corresponds to the jth row (i.e. j -th quantile) and t -th column of \mathbf{Y}_i in [Figure 3.](#page-5-0) A statistical model as in [Equation 1](#page-2-1) is then ftted to the summary statistics to obtain an estimated value $\hat{\theta}$. The statistical model we consider is [Equation 1](#page-2-1) with $f_i(\mathbf{x}_{it}, \boldsymbol{\theta}_i) = f_i(t, \Delta \rho_i, a_i, \boldsymbol{\theta}_i)$ and

$$
f_j(t, \Delta \rho_i, a_i, \mathbf{\Theta}_j) = \beta_{1j} + \beta_{2j}t + \beta_{3j}\Delta \rho_i + \beta_{4j}a_i
$$

+ $\beta_{5j}t\Delta \rho_i + \beta_{6j}ta_i + \beta_{7j}\Delta \rho_i a_i$
+ $\beta_{8j}t\Delta \rho_i a_i + \beta_{9j}t^2 a_i + \beta_{10j}t^3 a_i$
+ $\beta_{11j}t^4 a_i + \beta_{12j}t^2\Delta \rho_i a_i$
+ $\beta_{13j}t^3\Delta \rho_i a_i + \beta_{14}t^4\Delta \rho_i a_i.$ (3)

Therefore, the (unknown) parameter vector θ = $\begin{bmatrix} \mathbf{\theta}^T_1 \end{bmatrix}$ $\left[\mathbf{p}_1^T, \ldots, \mathbf{\theta}_J^T\right]^T$, with $\mathbf{\theta}_j = [\beta_{1j}, \ldots, \beta_{14j}, \sigma_j^2]^T$, has dimension $p = 15 \cdot J = 75$. To estimate the parameters in [Equation 3,](#page-4-1) namely $\hat{\theta}_j = [\hat{\beta}_{1j}, \ldots, \hat{\beta}_{14j}, \hat{\sigma}_j^2]^T$, we consider the Ordinary Least Squares (OLS) estimator for a standard

into J subsets and the corresponding Y_{iit} were regressed on $t, \Delta \rho_i, a_i$ and combinations of them according to [Equation 3,](#page-4-1) for each i separately. Importantly, in order to reduce the computational burden, we considered a uniform random subsample encompassing 30 % of the simulated interaction times t . Indeed, the total number of time points was, on average across the $m = 5$ simulations, approximately n_i = 2200, so that 30 % of the simulated interaction times represents an average of $n_i^* \approx 650$ points per simulation, which corresponds to a regression model with approximately $n_i^* \times m = 650 \times 5 = 3250$ responses. We remark that while the choice of summary statistics and statistical model is arbitrary, with the possibility of considering alternative options, the SE provides satisfactory results with the current formulation (see [Section 3.2\)](#page-5-1). Then, the (unknown) conditional expectation $\mathbb{E}\left[Y_{ij}\left|t, a_i, \Delta \rho_i, \boldsymbol{\theta}_j\right.\right] = f_j(t, \Delta \rho_i, a_i, \boldsymbol{\theta}_j)$ can be estimated by $f_i(t, \Delta \rho_i, a_i, \hat{\theta}_i)$. The latter can be used to predict t, or indeed also *a* or $\Delta \rho$, given a set of outcomes $y = [y_1, \ldots, y_J]^T$, which are usually obtained by feld measurements. In our case the measurements are the whole rock chemistry of erupted volcanic rocks. Thus, given an observed \hat{y} (field measurements of whole rock chemistry of erupted rocks), and as precise as possible information about a (from geophysics) and $\Delta \rho$ (which can be calculated from the chemistry of the two end-member magmas and existing models [see e.g. [Lange](#page-13-20) [1997\]](#page-13-20)), a statistical emulation prediction for the interaction time t can be obtained through

linear regression model. More precisely, the data were split

$$
\hat{t} = \underset{t}{\arg\min} \left\| \hat{y} - f(t, a, \Delta \rho, \hat{\boldsymbol{\theta}}) \right\|_{\hat{\Omega}(\hat{\boldsymbol{\theta}})}^2, \tag{4}
$$

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Figure 3: Flow chart from the simulator output to the statistical simulator and the prediction computation for the magma dynamics model in [Section 2.](#page-1-0)

where $\|\mathbf{x}\|_A^2 = \mathbf{x}^T A \mathbf{x}$, and $\hat{\mathbf{\Omega}}(\hat{\boldsymbol{\theta}})$ is a suitably chosen positive definite weighting matrix. In particular, $\hat{\Omega}(\hat{\theta})$ should be as close as possible to the inverse covariance matrix of the model's errors. In our case study we simply choose $\hat{\Omega}(\hat{\theta}) = \text{diag}(1/\hat{\sigma}_{p_1}^2, \ldots, 1/\hat{\sigma}_{p_J}^2)$, where diag(\cdot) denotes a diagonal matrix and $\hat{\sigma}_{p_i}$ (for $j = 1, \ldots, J$) is the standard error of the predicted mean for the j-th quantile. The prediction \hat{t} defned in [Equation 4](#page-4-2) is a classical minimum distance estimator with well-known statistical properties [see e.g. [Newey and](#page-13-23) [McFadden 1994\]](#page-13-23). Under plausible conditions, this estimator is consistent for t and, under stronger conditions, asymptotically normally distributed, which implies that we can use this property to associate prediction errors when assessing error propagation (see [Section 4\)](#page-6-0).

3.2 Prediction Accuracy

To evaluate the prediction accuracy of the SE, we consider \hat{y} which is obtained by running the simulator at known a , $\Delta \rho$, and t . Then, we can compare the predicted interaction time \hat{t} based on [Equation 4](#page-4-2) to the known interaction time t , say the *true* interaction time, hence providing a way to evaluate the (in sample) prediction accuracy.

[Figure 4A](#page-6-1) compares predicted (\hat{t}) versus the true (t) interaction time (both measured in hours) for Simulations 2–4 of [Table 1;](#page-2-0) Simulations 1 and 5 exhibit comparable results and are consequently excluded. The resulting predictions across

diferent simulation settings are very close to the true ones, which are represented by a dashed black line. This is further highlighted in [Figure 4B](#page-6-1) which reports the associated absolute prediction errors, i.e. $|\hat{t} - t|$. Overall, absolute errors are lower than 0.8 hours across all settings and tupically lower than 0.2 hours. However, a noticeable increase in prediction errors is present only for Simulation 2 as interaction time approaches 5 hours. This might be due to the fact that the setting of Simulation 2 produces outputs that take much longer than the others to homogenise, as both lower density contrasts and prolate reservoir shape hinder mingling efficiency [\[Montagna et al.](#page-13-16) [2015\]](#page-13-16), so that, at larger times, the output is notably diferent to the output in the other settings. This feature should be introduced in the SE, in order to produce better predictions, but this is left for further research. This prediction accuracy exercise highlights the subtle adjustments that are needed when comparing predictions to actual data. Moreover, it should be emphasised that larger prediction errors could be partially explained also by poorer performance of the SE itself in some parts of the parameter space (see [Section 5.2](#page-9-0) for further details).

Since the SE is built up using all information in the shallow reservoir, we also performed an emulation study to investigate two aspects of the SE; here we keep the SE parameters as estimated in [Section 3.1.](#page-3-2) First, we note that the whole simulated data for the estimation of the parameters of the SE (i.e. $\hat{\theta}_i$)

Figure 4: Panel [A] predicted (solid lines) versus true (dashed line) interaction times (in hours) across Simulations 2–4 of [Table 1,](#page-2-0) computed using [Equation 4.](#page-4-2) Panel [B] absolute prediction errors associated with each simulation setting.

for $j = 1, \ldots, J$ might not, *a priori*, be necessary. Therefore, we also considered uniformly subsampling the simulator's outputs for given aspect ratios, density diferences, and time, i.e. the rows of D_i in [Figure 3.](#page-5-0) The subsamples represent respectively 60, 30, and 10 % of the simulated outcomes, on which the quantiles of interest are then computed. Additionally, as an eruption could extract magma from progressively deeper portions of the magmatic reservoir, we assessed the accuracy of the SE in predicting the times considering the lower and upper half of the magma reservoir, separately. The predicted versus true interaction times (in hours) across Simulations 2–4, computed using [Equation 4,](#page-4-2) are presented in [Figure 5,](#page-7-0) for the diferent emulation settings. Panels [5A–5C](#page-7-0) assess the accuracy of the SE on the whole shallow reservoir, while panels [5D–5F](#page-7-0) and panels [5G–5I](#page-7-0) highlight its accuracy in respectively, the upper half and lower half parts of the shallow reservoir. Within each panel, the accuracy is assessed using diferent sampling proportions. In particular, [Figure 5A–C](#page-7-0) shows that, across the considered simulation settings, subsampling from the whole shallow reservoir does not signifcantly impact predictive accuracy, with the exception of Simulation 2 [\(Figure 5A\)](#page-7-0), which is characterised by a slightly longer homogenisation time, resulting in a more spread distribution. However, the remaining panels of [Figure 5](#page-7-0) highlight the presence of estimation biases even under a full sampling scheme. Such biases are more marked for Simulation 4, reported in [Figure 5C,](#page-7-0) [5F,](#page-7-0) and [5H,](#page-7-0) and they can be partially explained by the nature of this simulation. These biases can be generally associated with chamber stratifcation. Indeed, while the whole system is more or less going towards homogeneity, it does so in a way that the upper half shallow chamber contains more of one of the two end-member magmas, and the lower half shallow chamber contains more of the other. Having said that, it should be stressed that these biases are not due to an inaccurate adjustment of the SE, but instead

constitute additional information for feld data collection. For instance, if the available deposits from eruptions show a transition of chemistry from the bottom to the top of the deposit, then one would adapt the SE to separate the simulator outputs from the upper and lower half, so that better predictions can be built. In other words, when chemical gradients in erupted deposits can provide additional information on the modality of magma extraction from the subvolcanic reservoir, the SE can be adapted accordingly. Indeed, due to the non-homogeneous nature of reservoir processes, diferent regions of the system may record diferent homogenisation timescales, as opposed to one common history for all.

4 ERROR PROPAGATION

The (fixed) parameters a , $\Delta \rho$ and \hat{y} in [Equation 4](#page-4-2) are subject to measurement error, $f(\cdot)$ is subject to approximation error, and the (fixed) parameters $\hat{\theta}$ are subject to estimation error. The last two types of errors can be deduced from the properties of the estimators using model [Equation 2.](#page-3-1) We however found out that the approximation error—due to the use of a statistical model $f(\cdot)$ to approximate the true physical mechanism—is small for the considered application (see [Fig](#page-6-1)[ure 4\)](#page-6-1), and the estimation error for $\hat{\theta}$ is negligible.

Measurement errors for a , $\Delta \rho$, and the elements of \hat{y} can be inferred from scientists' experience. All sources of error can then be included in the statistical emulation procedure to obtain a set of possible values for \hat{t} in [Equation 4,](#page-4-2) from which confdence intervals can be built. This can be achieved through simulations. Indeed, one can consider that the errors are symmetrically distributed around the actual value of the parameters, so that, for each parameter, we can simulate an error drawn from a pre-specifed distribution, such as the normal distribution. This is equivalent to simulating possible values for the diferent parameters, from a normal distribution,

Figure 5: Predicted (solid lines) versus true (dashed lines) interaction times (in hours) across Simulations 2–4 of [Table 1,](#page-2-0) computed using [Equation 4.](#page-4-2) The predicted interaction times are computed on the whole shallow reservoir (panels [A]–[C]), the upper half part (panels $[D]$ –[F]) and the lower half part (panels $[G]$ –[I]) of the shallow reservoir. Within each panel, the predicted interaction times are computed on different (uniform) subsamples.

centred at the true value, with a pre-specifed variance. For the latter, based on domain knowledge, a standard error of 10 % of the actual value is usually expected. Let the elements $\hat{y}_{h,i}^*$ of \hat{y}_h^*, a_h^* , and $\Delta \rho_h^*$, for $h = 1, \ldots, H$, be (independent) random realisations drawn from normal distributions, then

$$
\hat{t}_h^* = \arg\min_t \left\| \hat{\mathbf{y}}_h^* - \hat{f}(t, a_h^*, \Delta \mathbf{p}_h^*, \hat{\boldsymbol{\theta}}) \right\|_{\hat{\Omega}(\hat{\boldsymbol{\theta}})}^2 \tag{5}
$$

represents a suitable distribution for \hat{t} that takes into account all (signifcant) sources of errors. Thus, a confdence interval at the 1 – α confidence level for *t* is given by $(\hat{t}_{(l)}^*, \hat{t}_{(u)}^*)$, where $l = |(\alpha/2)H|$ and $u = [(1 - \alpha/2)H]$, with $x_{(i)}$ being the *i*-th order statistics, and with $\lfloor x \rfloor$, respectively $\lceil x \rceil$, being the

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integer smaller than or equal, respectively larger than or equal, to x .

[Figure 6](#page-8-0) presents 95 % confdence intervals for predicted interaction times (in hours), across Simulations 2–4 (from left to right), when some parameters are fxed and others are subject to random error. Specifically, in Figure $6A-6C$, both a and $\Delta \rho$ are subject to random error while the elements of \hat{y} are fixed, in [Figure 6D–6F,](#page-8-0) a and $\Delta \rho$ are fixed while the elements of \hat{y} are subject to random error, and in [Figure 6G–6I](#page-8-0) all parameters are subject to random error. Overall, predicted interaction times are close to the true ones across all simulation settings, and we note that the estimation variability increases as we assess for all input parameters [\(Figure 6G–6I\)](#page-8-0). Confdence intervals are generally smaller than 20 %, reach-

Figure 6: Sensitivity of the SE across Simulations 2–4 of [Table 1,](#page-2-0) where input parameters $\Delta\rho^*$ and a^* (panels [A]–[C]), \hat{y}_i^* (panels [D]–[F]), $\Delta\rho^*$, a^* , and \hat{y}^*_j (panels [G]–[I]) are drawn from normal distributions centred at their true values $(\Delta\rho,a,\hat{y_j})$ with standard deviations equal to 10 % of such true values. Average predicted (solid lines) versus true (dashed lines) interaction times (in hours), as well as 95 % confidence intervals, are computed using *H* = 200 in [Equation 5.](#page-7-1) Each setting is restricted to a (uniform) subsample encompassing 25 % of the locations sampled from the shallow reservoir.

ing up to 30 %. Given the large uncertainties that characterise our understanding of inaccessible magmatic plumbing systems [\[Sigurdsson et al. 2015\]](#page-14-23), these confdence intervals are rather narrow and allow us to provide meaningful estimates of pre-eruptive magmatic interaction times.

5 DISCUSSION

The results of our analysis show that the SE is capable of reproducing the results of the numerical model with small uncertainty when both the contrast in density and the shape of the reservoir in which the interaction between magma occurs are known [\(Figure 4\)](#page-6-1). This is also true when we consider subsamples of the simulated domain. Interestingly, for some of the simulations, diferences emerge when considering the upper and lower half of the simulated domain [\(Figure 4E,](#page-6-1) [4F,](#page-6-1) [4H,](#page-6-1) [4I\)](#page-6-1). This indicates that for the specifc input parameters of these runs [\(Table 1\)](#page-2-0), the interaction will generate a zoned magma reservoir with chemical diferences between its lower and upper portions. This is of interest, as it shows the range of model input parameters under which the interaction between two magmas leads to chemical heterogeneities that can be identifed in the deposits of the eruption (e.g. chemical difference along the stratigraphy of deposit).

The contrast in density between two magmas can be calculated from existing models providing magma density as a function of its chemistry and the pressure and temperature conditions at which the interaction was occurring [e.g. [Lange](#page-13-20) [1997\]](#page-13-20)). The temperature and pressure at which the interaction occurs can be estimated using thermobarometry [\[Nimis](#page-13-24) [and Ulmer 1998;](#page-13-24) [Neave and Putirka 2017;](#page-13-25) [Petrelli et al. 2020;](#page-14-4) [Jorgenson et al. 2022\]](#page-12-23). Clearly, temperature and pressure estimates, together with estimates of the volatile content of magmas, which impact their density, are all associated with uncertainties. Additionally, the geometry of the reservoir could be defned using geophysics, but this would not be possible when studying past eruptions. Thus, to quantify the duration of the pre-eruptive interaction between magmas, and fundamentally, the associated uncertainty, it is essential to propagate all uncertainties associated with the input parameters of the model. This is when the advantages of statistical emulation become evident as the alternative would be to perform a large number of time-consuming numerical simulations varying the input parameters within a range considered reasonable by the experts. The results of runs performed using the SE and considering a normal distribution with 10 % uncertainty for input parameters such as α and $\Delta \rho$ and the measurable parameters in \hat{y} (i.e. the distribution of chemistry of erupted magma), show that the uncertainty in the estimated interaction time preceding an eruption increases with the duration of interaction, but does not exceed 25–30 % of the total interaction time [\(Figure 6\)](#page-8-0).

As a byproduct, the SE provides insights on the sensitivity of model results to specifc input parameters. Successive runs of the simulator must thus target specifc regions of the parameter space that yield most representative results. Moreover, the SE results obtained with large undersampling of the simulated outputs are practically indistinguishable from those obtained with the whole datasets [\(Figure 4\)](#page-6-1). Eliminating unnecessary outputs would largely decrease computational costs of large simulations, contributing to an overall leaner and more manageable workflow.

5.1 Application to petrological dataset

To test the validity of the SE, we have performed a proof-ofconcept inversion on a petrological dataset from the Agnano-Monte Spina eruption at Campi Flegrei, on which the numerical model is loosely based [\[Arienzo et al. 2010\]](#page-11-14). The products of this eruption were selected because of existing evidence that magma mixing ultimately triggered this event and it is considered as the reference for future large-size explosive eruptions at Campi Flegrei [\[D'Antonio et al. 1999;](#page-11-17) [Di Vito et al. 1999\]](#page-12-24). We used Al_2O_3 weight fraction in melt inclusions as a proxy for composition C . We express composition as weighted sum of phonolitic and shoshonitic (indexed by P and S, respectively) end-members: $C = \alpha C_P + (1 - \alpha C_S)$, where α is phonolite weight fraction. The compositional distribution obtained from the dataset is then fed into the SE to obtain the pre-eruptive interaction times, given a distribution of the two parameters a and $Δρ$. The distributions for a and $Δρ$ have been chosen, respectively, as skewed towards sill-like geometries and uniform. The obtained interaction times range between 5 and 8 hours [\(Figure 7\)](#page-10-0), which is a timescale loosely compatible with what has been proposed for the recharge-to-eruption times at Campi Flegrei (hours to days) [\[Perugini et al. 2010;](#page-14-24) [Arzilli et al. 2016;](#page-11-18) [Astbury et al. 2018\]](#page-11-19). In the simulations we consider all analyses collected along the stratigraphy together, which return a unimodal distribution [\(Figure 7B\)](#page-10-0). Some chemical zoning was present in the reservoir as testifed by the increase of mafc component toward the top of the stratigraphic sequence [\[Arienzo et al. 2010\]](#page-11-14). However, analyses collected for the same eruptive unit would not be sufficient to produce a statistically representative distribution of whole rock compositions to compare with the SE. Because with the progression of mixing, the compositional distribution of magma in the reservoir evolves from bimodal to unimodal, our results tend to overestimate the timescale of mixing. These considerations provide a framework for the collection of samples on which to apply the approach we propose. More precise constraints of the timescales of mixing would require the collection of a statistically representative number of samples for each eruptive unit and possibly on the largest possible number of outcrops at diferent locations. Regarding the case at hand, the extremely short timescale of magma interaction that preceded the Agnano-Monte Spina eruption provides solid and important constraints that can help in assessing volcanic hazards.

5.2 Limitations and further development

Some limitations of our modelling approach and potential avenues for further research, as highlighted by one of the anonymous reviewers, deserve additional consideration. Firstly, [Fig](#page-6-1)[ure 4](#page-6-1) highlights the potential presence of some systematic biases for the considered SE, particularly in the range between 2 and 3.2 hours, where predicted interaction times consistently overestimate the true values. It also suggests that the model may not be fully reliable at the boundary (e.g. at larger interaction times) where predicted interaction times can be sensitive to slight changes in the true time. These discrepancies may arise from several sources. In particular, the (function of the) simulated data lie in the interval (0, 1); we use a regression model for data in (−∞, ∞), meaning that a beta regression would be more appropriate, for example. However, in most settings of practical relevance, we obtained predictions (as well as their associated prediction intervals) between 0 and 1 (results not shown here). Another source of inaccuracy could be attributed to unaccounted correlations in the simulated data that can result into unaccounted correlations in the model's residuals. In the settings of practical relevance, we found that the residuals from our linear model are fairly well-behaved, although they exhibit some "mild" form of longrange dependency (results not shown here). Additionally and in general, the SE should also be adapted to the potential presence of non-linear relationships and heteroskedasticity. Moreover, the ftting method should also be adapted in the cases where outlying observations might be present.

We acknowledge the risk of model misspecification in general, which for the cases studied in this paper, appear to be quite mild. Indeed, as illustrated in [Figure 4B,](#page-6-1) the prediction error is quite small for practical purposes. Moreover, we remark that simpler models, such as the one in use, often have a tendency to generalise better than more complex methods. While our approach maintains consistency against misspecifed models, future research may explore more suitable and

Figure 7: [A] magmatic interaction times as obtained from the SE inversion of a compositional dataset from products of the Agnano-Monte Spina eruption at Campi Flegrei across a grid of a and $\Delta\rho$ inputs. Interaction times recovered from the model match those provided in the literature for the specific eruption. [B] distribution of predicted interaction times (in hours), as well as 95 % confidence intervals and median (dashed vertical lines) based on $H = 10,000$ replications in [Equation 5.](#page-7-1) The sensitivity of the SE was assessed at random input parameters $\Delta\rho^*$ and a^* , which are respectively drawn from a uniform and a beta distribution with shape parameters α = 2 and β = 5 (represented as light green and blue densities in panel [A], respectively) and \hat{y}_j^* which is drawn from a normal distribution centred at its true values \hat{y}_i with standard deviations equal to 10 % of such true values.

fexible approaches such as semi-parametric methods. In our modelling exercise, we mainly focused on point estimation and, as highlighted in [Figure 6,](#page-8-0) uncertainties in $\hat{\theta}$ are dominated by input-related factors. The overall performance of the considered SE is satisfactory for the considered case study, leading to good predictive power and generalisability.

We also remark that a very limited input grid was used in our experiments, and we are currently exploring the use of a fner grid to achieve better results. Overall, our study serves as a proof-of-concept, recognising limitations and paving the way for future enhancements in methodology and data collection.

6 CONCLUSIONS

The Earth Sciences often focus on phenomena that occurred in the past or at inaccessible depths, which inhibits our capacity to defne with confdence the range of input parameters of numerical models targeting the quantifcation of geological processes. Additionally, because of the complexity of some physical models, the realisation of a large number of simulations which would allow one to assess how uncertainty in the input parameter afects the fnal results—is prohibitive. Here, we use a specifc application to volcanology that allows us to highlight the multiple advantages of combining physical modelling and statistical emulation. We show how to determine the duration of magma interaction preceding a volcanic eruption from the analysis of the distribution of the erupted rock chemistry. This particular case is appropriate for Campi Flegrei (Italy) [\[Bagagli](#page-11-7) [et al. 2017;](#page-11-7) [Morgavi et al. 2017;](#page-13-15) [Montagna et al. 2022\]](#page-13-12). The approach we present here could be applied to past eruptions of Campi Flegrei to determine the duration of magma interaction preceding any volcanic eruption throughout its eruptive history. The workflow would require:

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i. The analysis of the whole rock chemistry of a statistically representative set of samples of each eruption.

ii. The calibration of SE on physical models realised over a wide range of geometries, magma conditions and duration.

iii. The comparison between the measured and calculated distribution.

The results of our analysis show that statistical emulation provides an excellent means to quantify the uncertainty in duration of magma interaction preceding a volcanic eruption from the distribution of chemical composition of the erupted rocks.

AUTHOR CONTRIBUTIONS

All authors contributed to the conceptualisation and methodology of the paper. CPM provided the raw simulated data from physical modelling. LI and SG developed the ^R code and performed the statistical analyses. All authors contributed to the writing and editing of the manuscript.

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DATA AVAILABILITY

The data analysed in this paper, as well as documented ^R code to process the data and replicate our results, are openly available at: <https://github.com/LucaIns/magma>. The code used for the magma dynamics simulations is available at: <https://gitlab.com/dgmaths9/gales>.

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